

# IMEX – Interplanetary Meteoroid Environment for exploration:

Modeling meteor showers anywhere, anytime

R. Soja, M. Sommer, R. Srama, P. Strub, E. Grün, J. Rodmann, J. Vaubaillon, A. Hornig, L. Bausch, J. Herzog



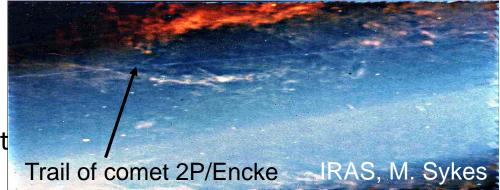


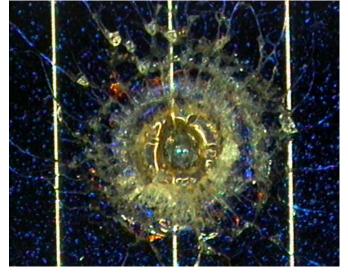
#### **Goal:**

Model of the fine structure in the Solar System dust cloud from cometary trails/streams (>100µm)

#### To study:

- meteor showers at Earth
- meteor showers at other planet locations
- comet dust emission
- planetary perturbations, scattering
- > the impact hazard to spacecraft





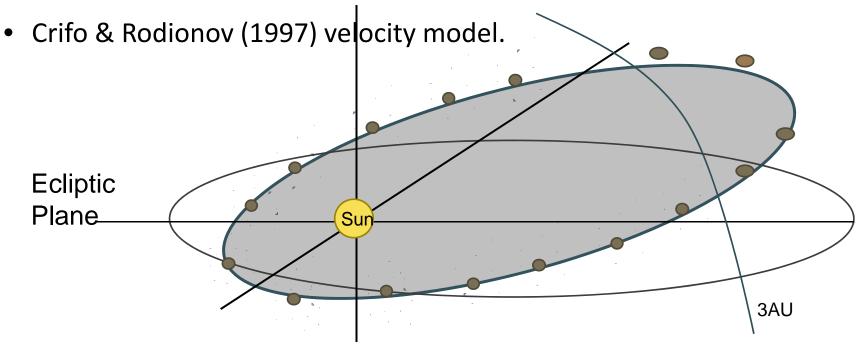
Impact on to Hubble

Space Telescope solar cells ESA, Drolshagen, 2008





- **Emit** dust:
  - from 428 short comets
  - for 200-400 years
  - inwards of 3AU from sunlight hemisphere
  - around 0.33 million particles/ comet/ mass
  - 8 masses 100microns to 1cm.



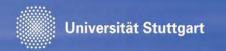


### (2) Creating the meteoroid stream database

- Integrate orbits (gravity and radiation effects)
  - Constellation distributed computing platform
  - Citizen science project: 12478 users with 57,521 PCs world wide
- Generate a database of cometary dust
  - Save each integrated particle several times per orbit between 1980-2080
- ➤ Result: we have a database from which can reconstruct the orbits of all stream particles from 429 comets.



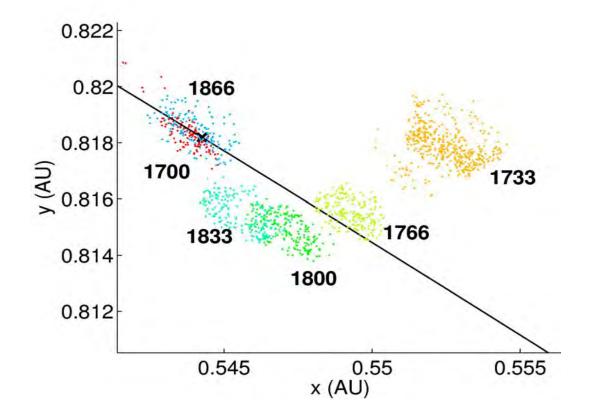
Andreas Hornig and Lars Bausch: aerospaceresearch.net





## Verifying the model (1): Meteor Storms Leonids in 2001

- Dust from **55P/Tempel-Tuttle** 1690-1998
  - Leonids at Earth 2001
  - Compare observed and modeled max times of storms at Earth: Within 10-30mins.

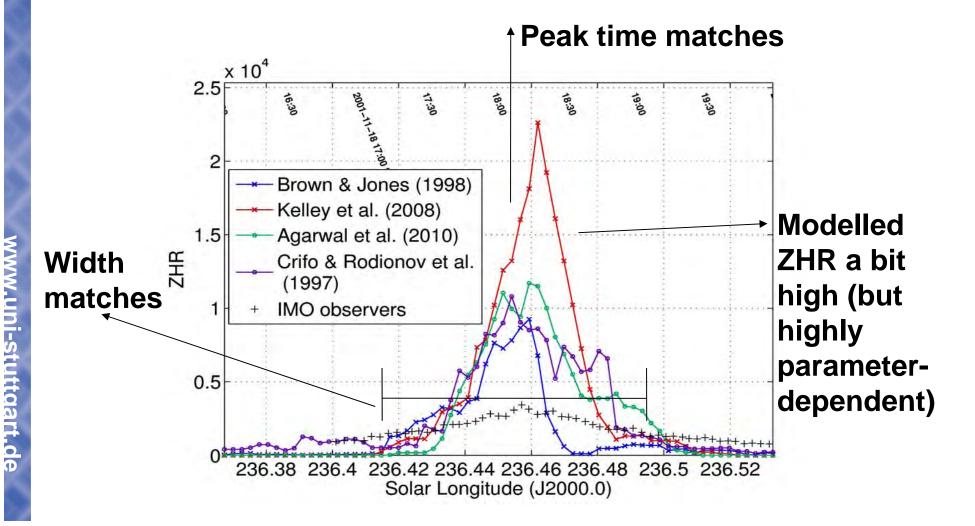






## Verifying the model (1): Meteor Storms Leonids in 2001

• ZHR profiles for different velocity models



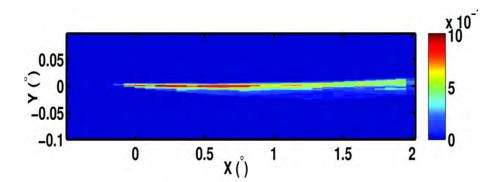




# Verifying the model (2): Comet Trails Comparison with IR images

- > 67P/Churyumov-Gerasimenko
  - Spitzer observations in 2004-6
    - Kelley et al 2008
    - Agarwal et al. 2010
  - Brightness match factor ~2





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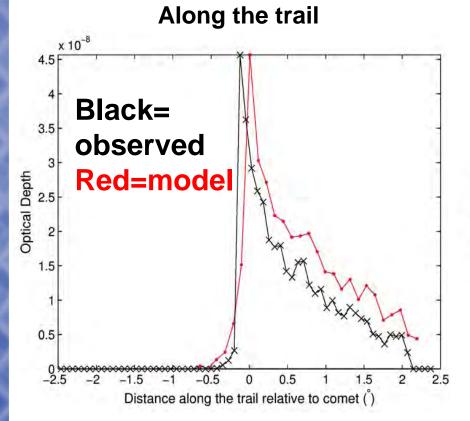




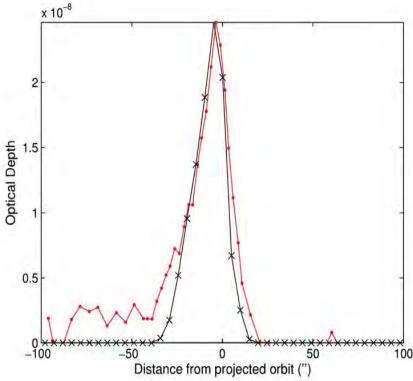
# Verifying the model (2): Comet Trails Comparison with IR images

Trail width matches well, but requires low velocities (Agarwal et al. 2010 model)

#### Kelley et al. 2008 observations: trail profiles



#### Across the trail







### **Summary**

- > IMEX model now performs well for individual streams
  - Generation of large volumes of data works well
  - Resulting trails match IR trail observations and meteor storm observations
- Ready for the next stage:
  - mass production of database of streams of 429 short period comets for the period 1950-2100.







### The Trail of 67P/Churyumov-Gerasimenko





### **Dust Streams near an object near 1AU (STEREO-B)**