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Automated detection and analysis of Moon impact flashes from Spain

José María Madiedo^{1,2}, José Luis Ortiz Moreno³, Nicolás Francisco Morales Palomino³ and Jesús Cabrera Caño²

¹ Facultad de Ciencias Experimentales, Universidad de Huelva, E-21071 Huelva, Spain

² Facultad de Física, Universidad de Sevilla, Departamento de Física Atómica, Molecular y Nuclear, E-41012 Sevilla, Spain madiedo@cica.es

³Instituto de Astrofísica de Andalucía, CSIC, Apt. 3004, Camino Bajo de Huetor 50, E-18080 Granada, Spain

We are currently performing a monitoring of the night side of the Moon in order to identify flashes produced by the impact of meteoroids on the lunar surface. For this purpose, we employ several telescopes equipped with high-sensitivity CCD video cameras. Software development plays an important role in our project, and, as a result of this, our detection and analysis package has been improved. Some of the results obtained so far are presented here.

1 Introduction

One of the techniques suitable for estimating the flux of interplanetary matter impacting the Earth is based on the monitoring of the night side of the Moon to detect flashes produced by the impact of meteoroids on the lunar surface. The first steps to detect such flashes date back to 1999 (Ortiz et al., 1999). Thus, impact flashes were unambiguously detected during several major meteor showers by using this technique and flashes of sporadic origin were also recorded (Ortiz et al., 2000; 2002; 2005; Cudnik et al., 2002; Cooke et al., 2006). In this context our team has developed the MIDAS project (Moon Impacts Detection and Analysis System). Its main objective is the identification of these flashes by means of an automated system that employs small telescopes operating from different locations in Spain. The identification of these events is performed by the MI-DAS software (Moon Impacts Detection and Analysis Software).

2 Methods and preliminary results

Figure 1 shows the location of our impact flashes detection systems. We currently monitor the Moon from Sevilla (Southern Spain). Our observatory can employ several SC Celestron telescopes (two C14, one C11, and one C9.25), all of them endowed with Watec 902H Ultimate CCD video cameras. In addition, a new 40-cm telescope is being setup in Central Spain, which will be operated from La Hita Astronomical Observatory.

GPS time inserters are used to stamp time on every video frame with a precision of 0.001 seconds. Meade 3.3 focal reducers are also used. With this configuration, we monitor about 5.8×10^6 km² on the lunar surface, give or take 10%. Sufficiently large lunar features are easily visible in the earthshine, and these can be



 $Figure \ 1$ – Location of the Moon impact flashes detection systems.

used to determine the selenographic coordinates of impact flashes. The Watec CCD cameras we employ work according to the PAL standard (25 fps), with a resolution of 720×576 pixels. The images taken by them are stored and digitized on multimedia hard disks. Then, they are sent to a computer for further processing and analysis.

The MIDAS software (Moon Impacts Detection and Analysis Software) was developed to automatically identify impact candidates (Madiedo, 2010). The main kernel in the software is related to the automated identification of impact flashes. Several algorithms are available to perform these detections, and new algorithms are currently being implemented in order to increase the efficiency of the software. During the detection process, a database with potential impact candidates is created for every telescope. These databases are automatically compared to establish which events are produced by the impact of meteoroids and which of them are related to other phenomena (cosmic rays, etc.). The software can also perform a photometric analysis of impact flashes in order to calculate the absolute magnitude of these events and the mass of the impactors. Figures 2 and 3 show confirmed impact flashes detected on March 27, 2012, at $20^{h}47^{m}16^{s}.281 \pm 0^{s}.001$ UT and July 26, 2012, at $21^{h}35^{m}04^{s}.686 \pm 0^{s}.001^{s}$ UT, respectively.

MIDAS is endowed with additional tools, such as video processing filters, a lunar phase calendar and a database containing information about meteoroid streams.

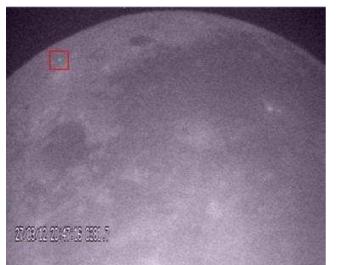


Figure 2 – Confirmed impact flash recorded from Sevilla on March 27, 2012, at $20^{h}47^{m}16$ ^s281 UT.

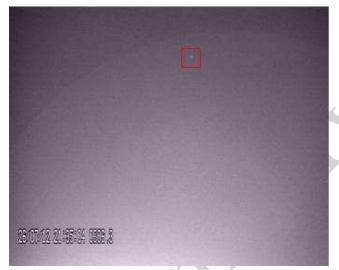


Figure 3 – Confirmed impact flash recorded from Sevilla on July 26, 2012, at $21^{\rm h}35^{\rm m}04\overset{\rm s}{.}686$ UT.

3 Conclusions and future work

We are operating a system that monitors the night side of the Moon in order to detect flashes produced by the collision of meteoroids on the lunar surface. A software package (MIDAS) has been developed to automatically identify and analyze these flashes. Our detection network will be expanded in a near future, as we are currently setting up a new telescope in Central Spain. On the other hand, new detection algorithms are currently being implemented in the MIDAS software. As a result of this we expect to identify impact flashes more efficiently.

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