

Evolution of Halley's Comet & Orionid Stream

(Supervised by Dr David Asher)

Aswin Sekhar

Ph.D Student

Armagh Observatory & Queen's University Belfast

United Kingdom

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Plan of Talk

1. Brief Introduction- Relevant Existing Works on Comet Halley & Resonance of Orionid Stream

2. Initial Results

3. The Final Goal

1. Relevant Existing Works:

1.1. Long Term Evolution and Life Time of Comet Halley

- * Observational records date back to 240 B.C. (Yeomans & Kiang 1981)
- * Accurate orbital calculations are available from 1404 B.C. (Yeomans & Kiang 1981)
- * Comet Halley's dynamical lifetime is estimated of the order of 100,000 years (Hughes 1985, Hadjuk 1986, Steel 1987, Emel'yanenko & Bailey 1996)

Photograph of Comet Halley (Source: W. Liller, IHW, 8 March 1986)



ESA's Giotto Image of Comet Halley



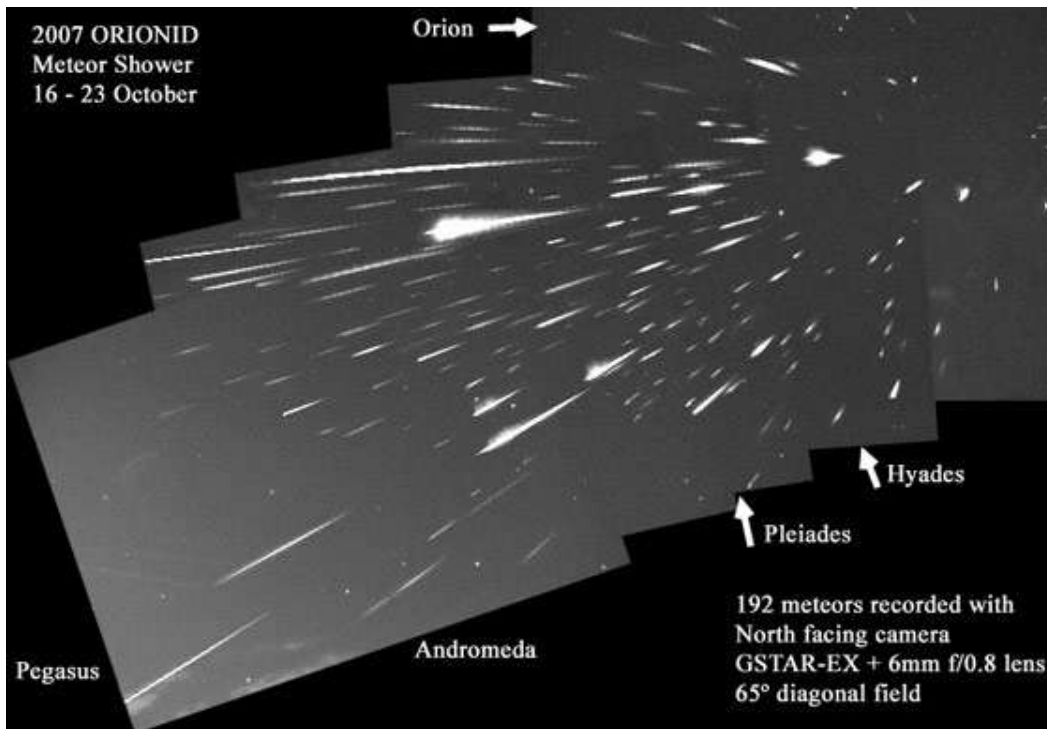
1.2. Resonance of Orionid Stream

* Peaking of Orionids in 2006 was unexpected

* Rendtel proposed 1:6 resonance as a mechanism (Rendtel 2007), noting also enhanced Orionid rates between 1933-1938 (Lovell 1954)

* It was also found that streams ejected in -1,265, -1,197 and -910 approached Earth in 2006 (Sato & Watanabe 2007)

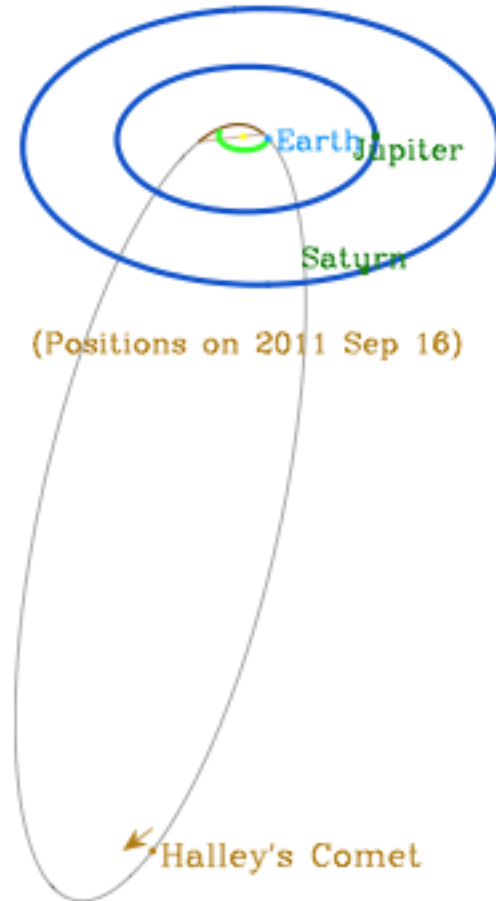
(Source: Astron.Assoc.Queensland)



Engraving of 1833 Leonids Storm



- Orbit of Comet Halley as of 16 Sep 2011



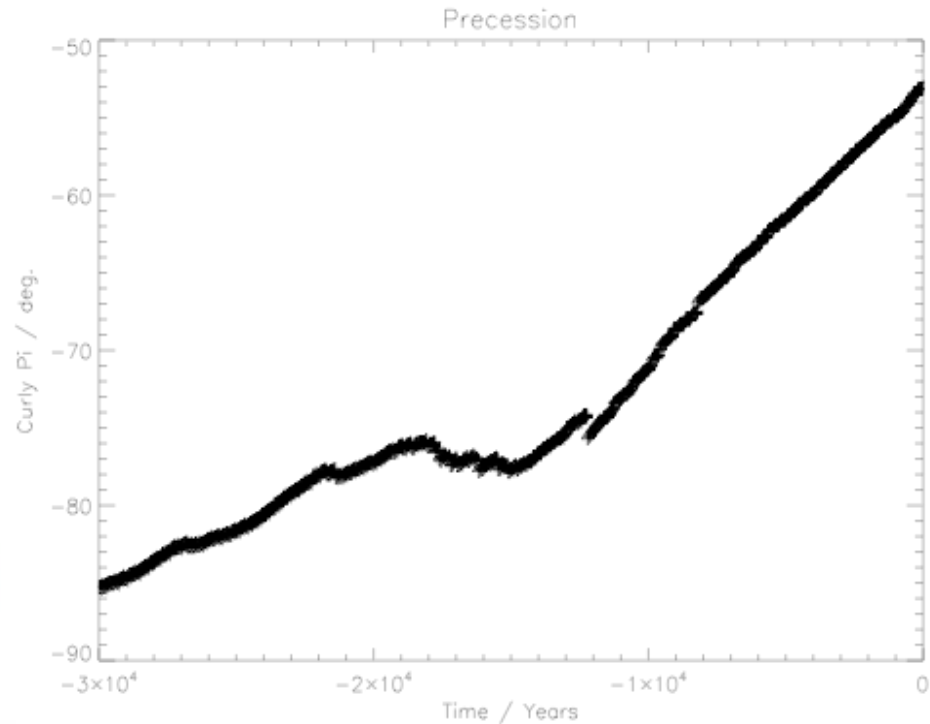
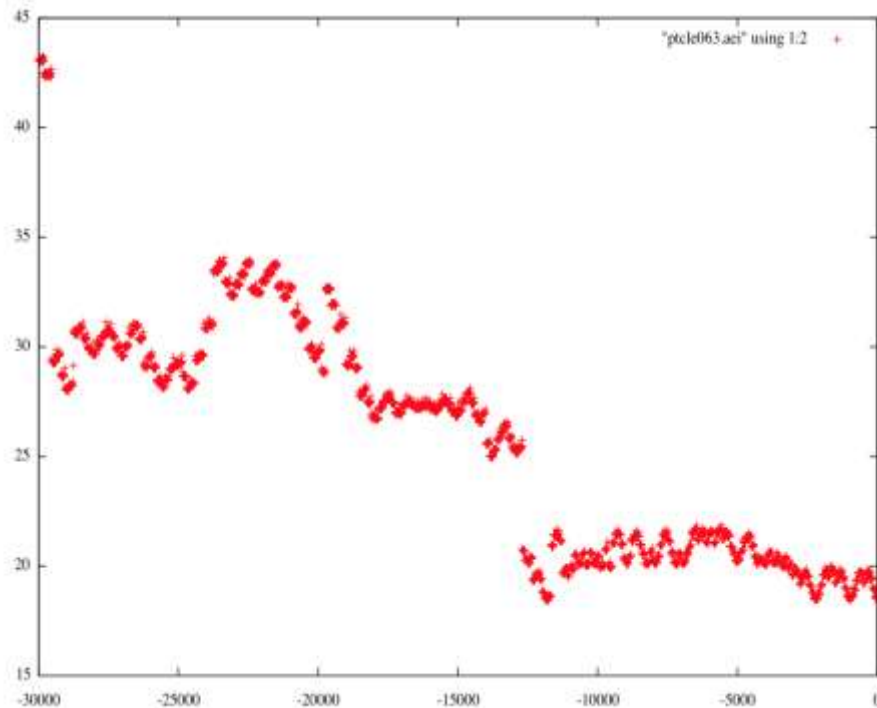
2. Initial Results

2.1. Comet Halley's Present Orbit

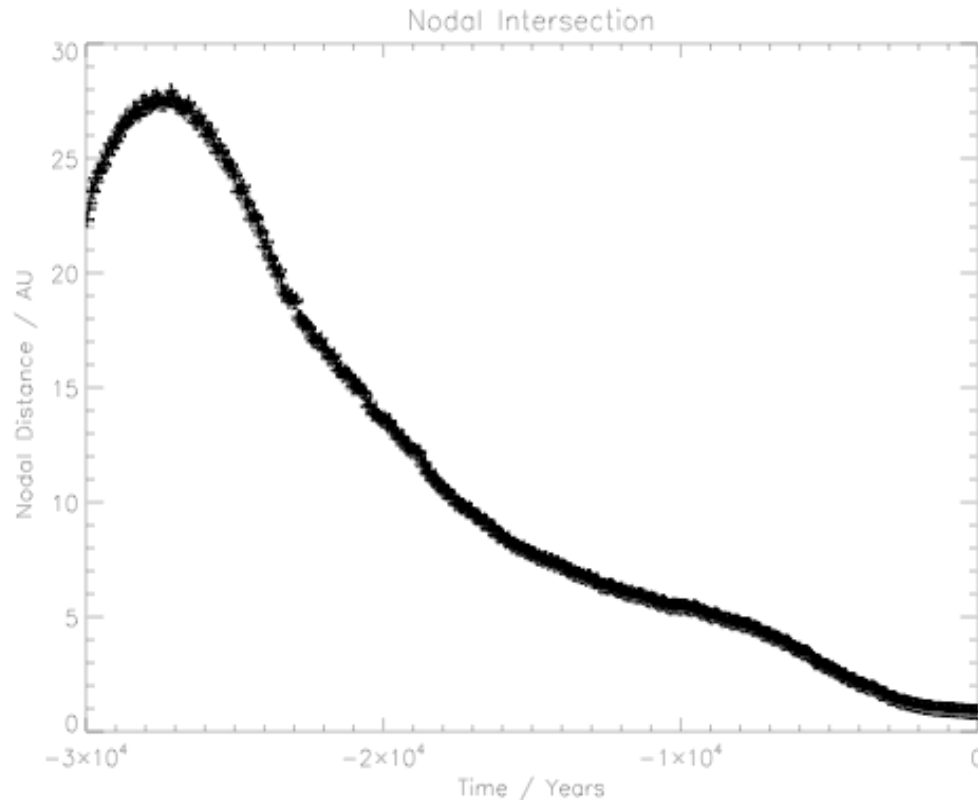
* Backward integration for 30,000 years was done for 100 test particles

* Longitude of pericentre (measure of precession of orbit) shows a significant variation (and also a discontinuity in many cases)

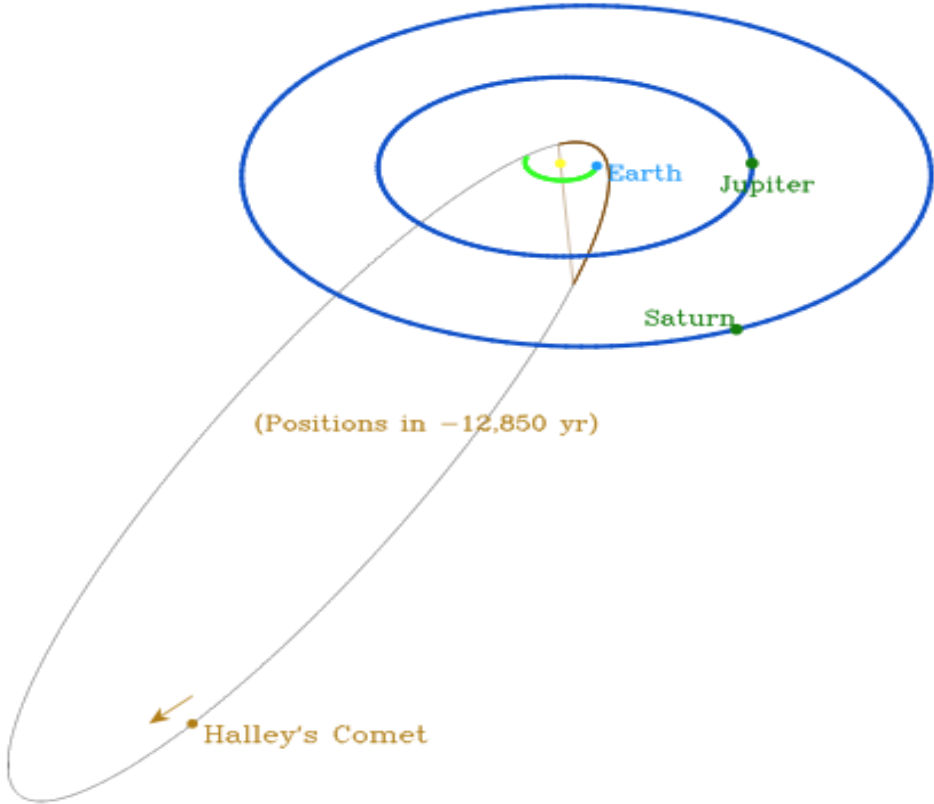
Semi-Major Axis (AU) Vs Time (Years)



- * The heliocentric distances for ascending and descending nodes were analysed.
- * It shows that there are possibilities for close encounters with Jupiter around this -12,000 time frame.
- * Hence we assume that most of the particles would have got strong perturbations from Jupiter which suddenly changed its orbital parameters.
- * Drastic variation in semi-major axis near -12,000. Orbits before this time cannot be very reliable for ejection models.

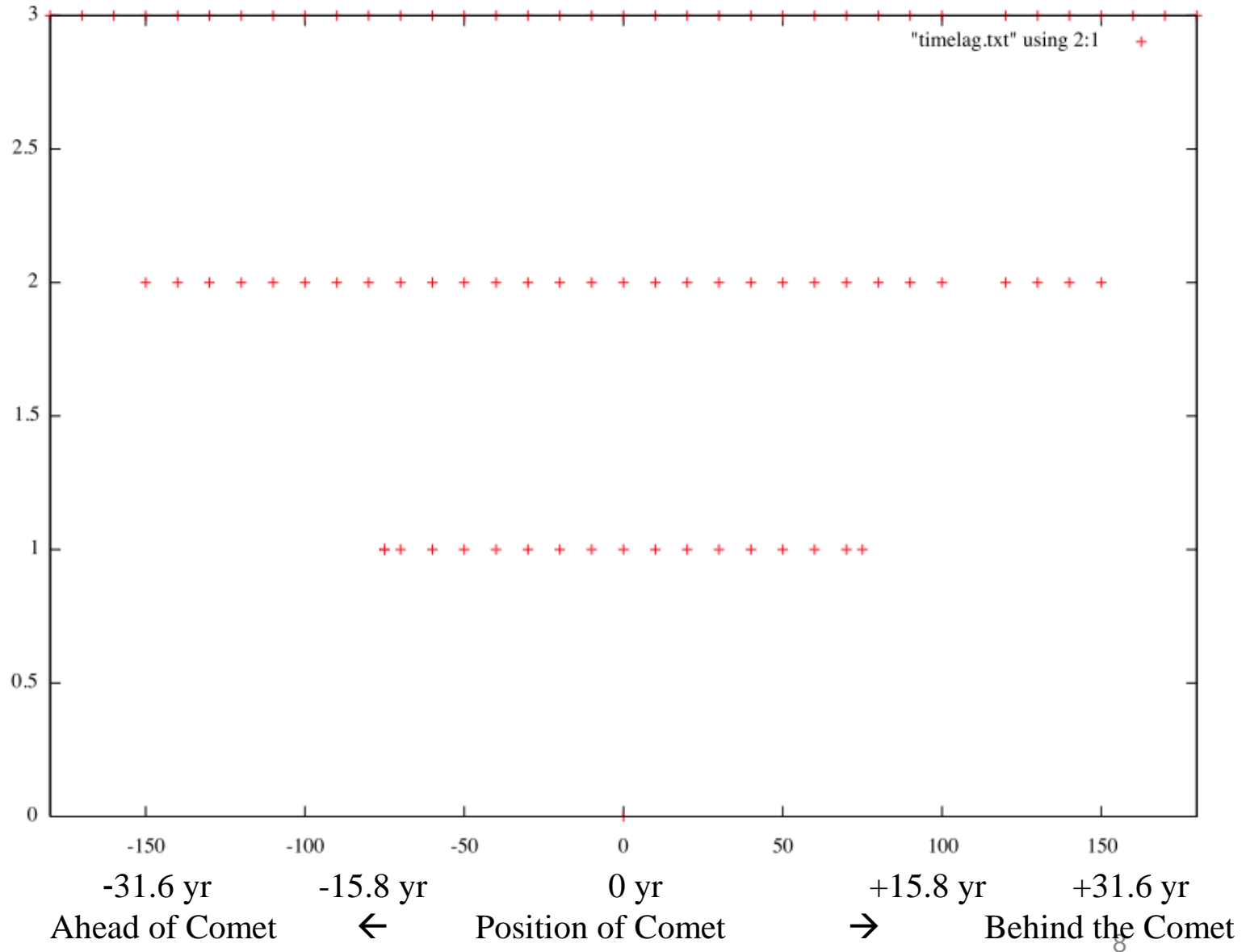


Orbit of Comet Halley in -12,850



2.2. Mean Anomaly Distribution

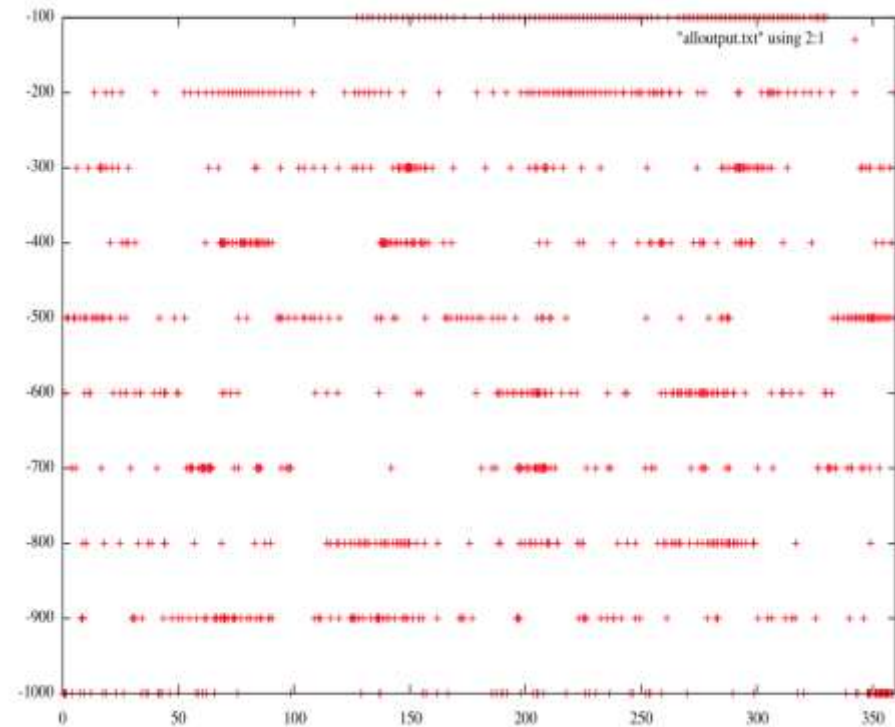
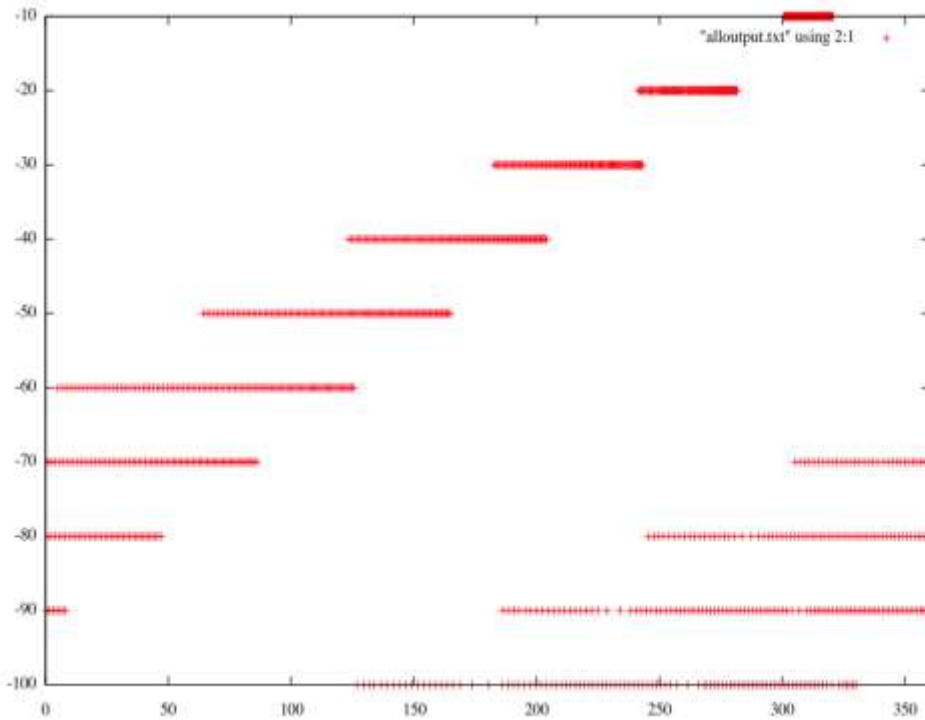
No. of Revolutions Vs Mean Anomaly (Degrees)



* Mean anomaly distribution versus time for 100 test particles shows that particles spread along the entire orbit in due course of time

*The distribution is roughly uniform

Time (Years) Vs Mean Anomaly (Degrees from 0-360)



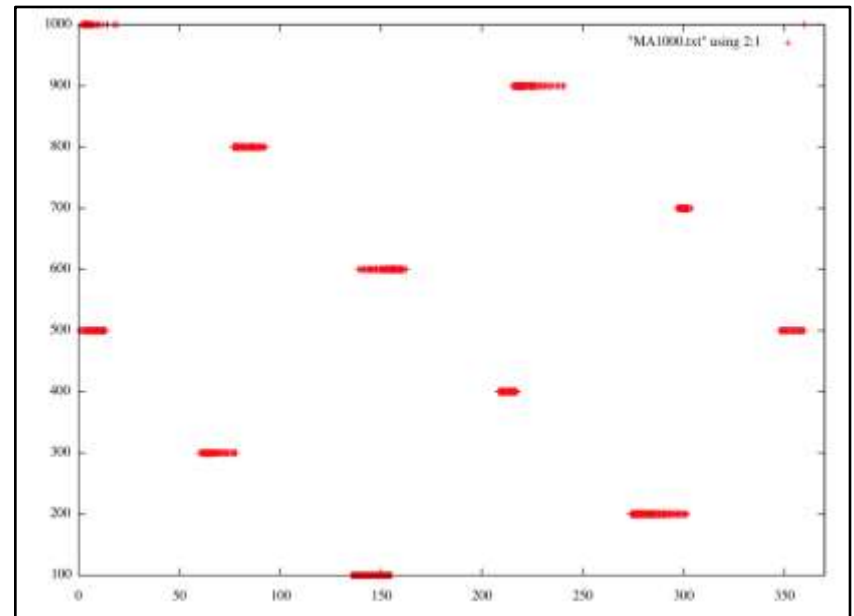
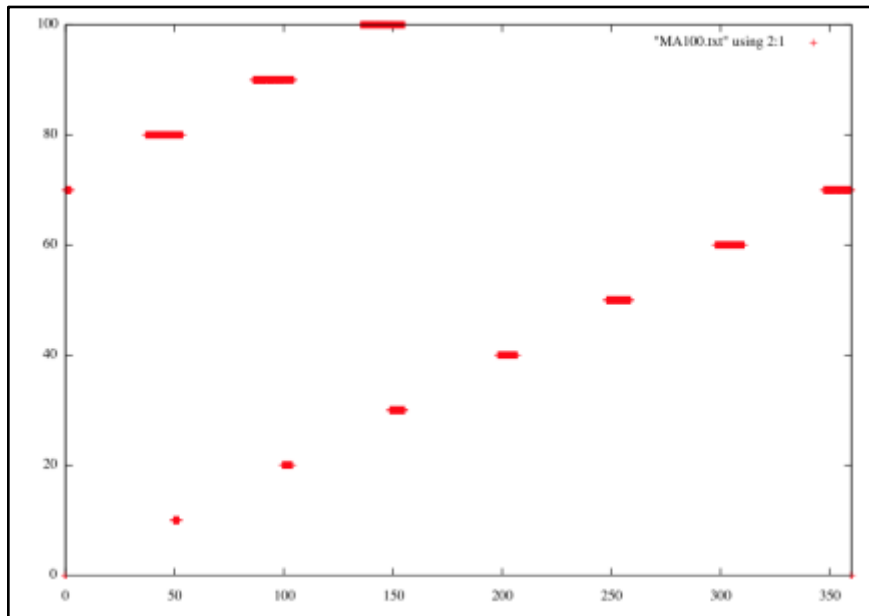
* Particles initially resonant behave differently

* They tend to stay together during its evolution because of periodic effects from Jupiter

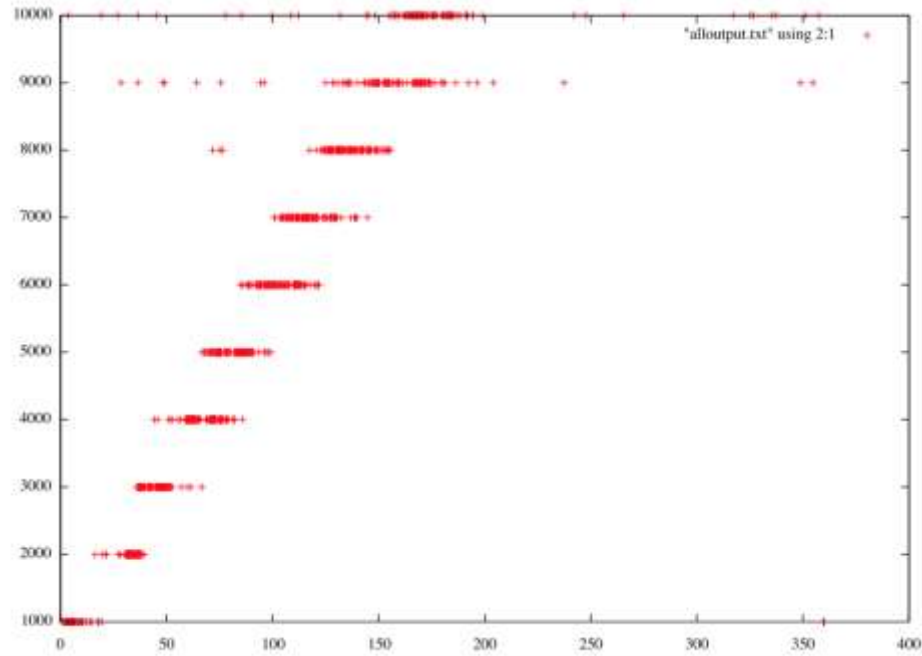
* Such clustering of particles in space leads to meteor outbursts

* Preliminary analysis shows about consecutive 5-6 years of enhanced Orionid activity

Time (Years) Vs Mean Anomaly (Degrees from 0-360)



Time (Years) Vs Mean Anomaly (Degrees from 0-360)



VALUABLE STATISTICS FROM IMO ZHR

Year	IMO	IAU Telegram
ORI 2010	40	40-50 predicted
ORI 2009	45	32 observed
ORI 2008	40	40 predicted
ORI 2007	70	65 +/- 3 observed
ORI 2006	>50	
ORI 1993		35 observed
ORI outbursts observed during 1933-1938, peaked in 1936 (Lovell 1954)		

- Our calculations show that 1:6 resonance can occur if semi-major axis of the comet lies between 16.6 AU to 17.8 AU
- Hence it is possible that Halley's comet itself was resonant in ancient times.
- An ejection model incorporating this idea can populate a lot of resonant meteoroids in 3D space & track them forward in time
- Quantifying various contributions from other resonances with Jupiter like 2:13, 1:7 etc is another interesting area

3. Final Goal

- 3.1. To try different ejection epochs and extend the present analysis of Halley stream to correlate with past observed events of meteor outbursts and predict such events in future
- 3.2. Develop a fundamental understanding of outgassing phenomenon and the dynamical evolution of such meteoroids over long periods of time by developing a 3D phase space ejection model.

ENOUGH TO KEEP US BUSY TILL 2061 (ON AN
OPTIMISTIC NOTE!!)

THANK YOU

