# The hard task of observing the meteoroids

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#### Introduction

- Too small to be detected by telescopes
- Too large to efficiently scatter Sun light
- 90% of cometary mass loss

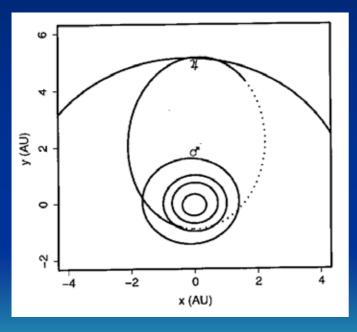






### Limits of today observations

 Meteors: necessary but not enough



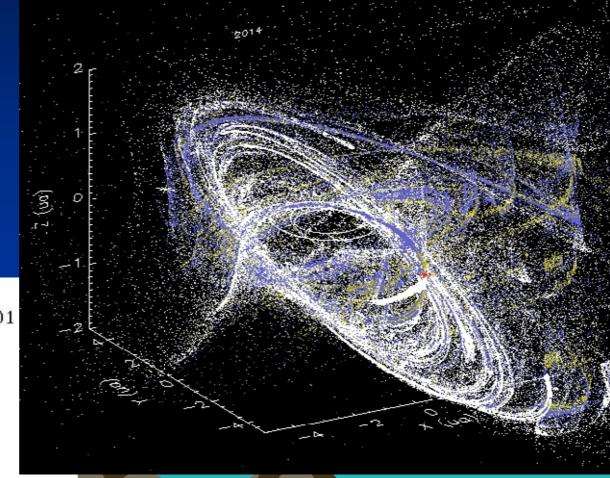
Wiegert et al. 2005, MNRAS, 361, 638-644

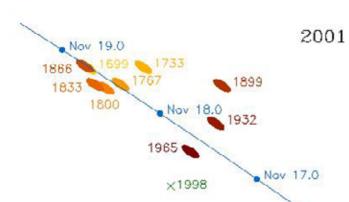


Future French network

= LORMS (Maquet, Vaubaillon, Colas, Vienne 2006)

# What we see and what we want to know (1)





D. Asher & R. McNaught (1999)

67P/Churyumov-Gerasimenko

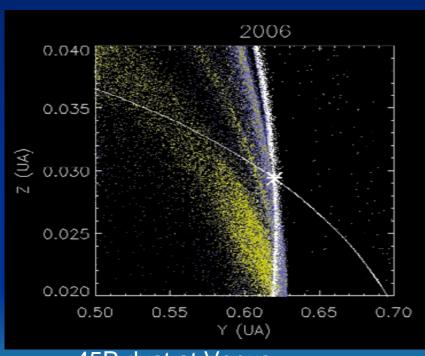
### Extend our reach (1)

- Other planets
- Christou et al., 2007

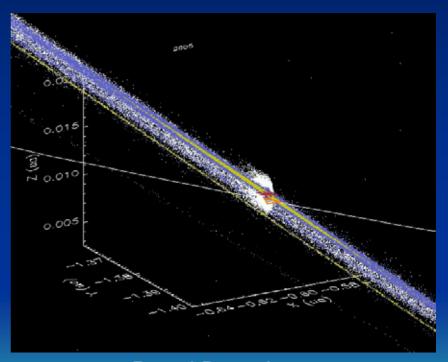
   (in press, Europlanet):
   review of tomorrow
   possibilities: visible,
   radio, exosphere etc.
- Dedicated instruments (ESA study, D. Koschny)



### Extend our reach (2)



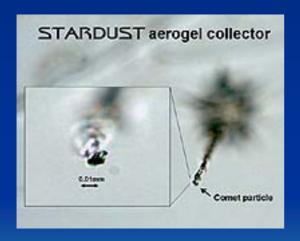
45P dust at Venus (Vaubaillon, Christou, 2006)



9P and Deep Impact (Vaubaillon et al., 2006)

#### In situ measurements

The question of the grain density

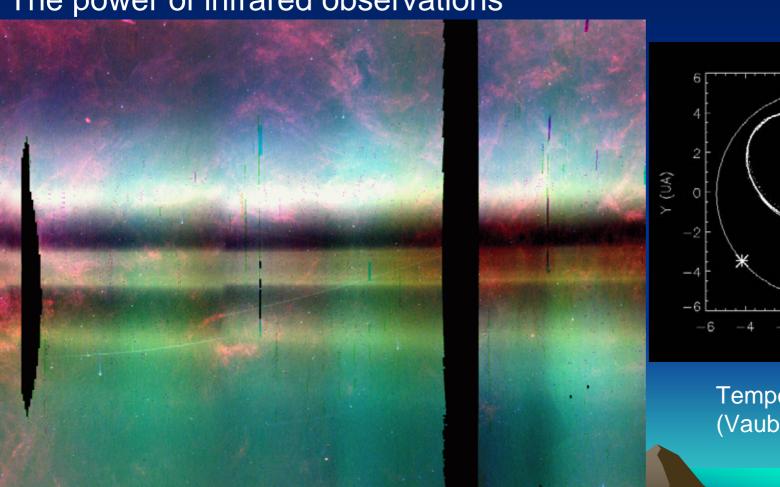


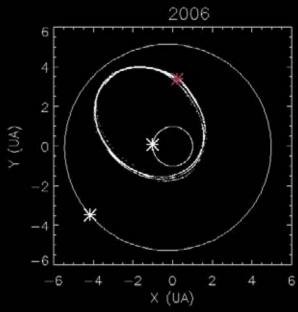




#### What we see and what we want to know (2)

The power of infrared observations





Tempel 2: simulations (Vaubaillon, Reach)

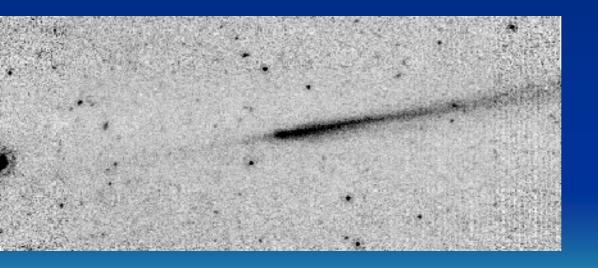
Tempel 2: IRAS observations

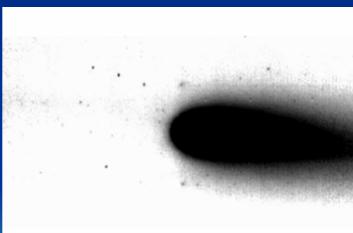
# Infrared observation (1)

• Reach et al., 2007: survey of 32 comets

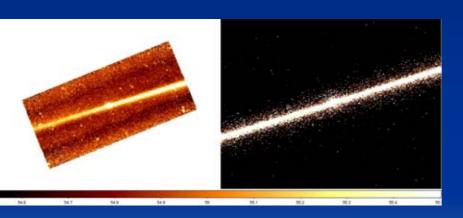
67P/Churyumov-Gerasimenko

88P/Howell





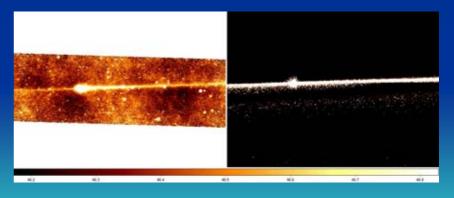
# Infrared observation (2)



10P/Tempel 2



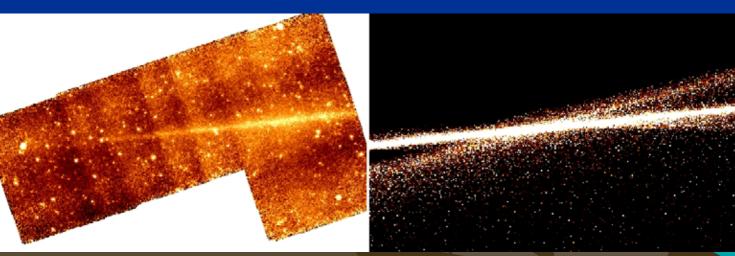
Clark



# Infrared observation (3)



Slaughter-Burnham



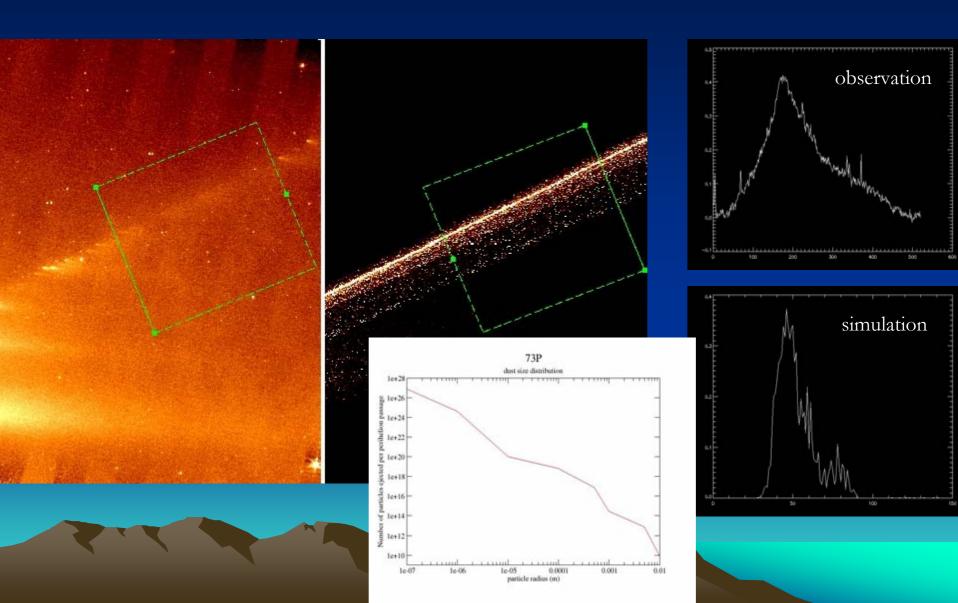
Clark (2005)

#### IR observations of 73P

- Age of particles
- Q\_dust before/after breakup ~1/11
- Pb with coma
- Puzzling spectrum



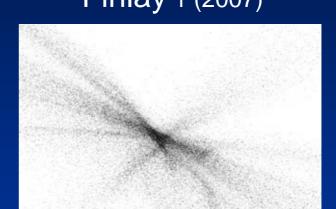
#### Size distribution index



# What we (well... I at least...) dream to see Finlay 1 (2007)

32P/Coma Sola (2013)





45P/Honda-Mrkos-Pajdusakova (2007)



#### Conclusion

- Meteor observation on Earth is necessary but not enough
- Other planets
- Infrared
- In situ

Need for an unaltered cometary dust sample return