# Integration of mean orbits of meteoroid streams 

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## Plan

(1) Traditional approach for mean orbit calculations
(2) Jopek et al. method
(3) Example of Neuschwanstein
(4) Example of Quadrantids

## Calculation of mean orbit

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Instead one can use: $1 / a, e, i, \omega, \Omega$.
or: $1 / a, q, i, \omega, \Omega$. even: $q, Q, i, \omega, \Omega$ can be used.

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- Meteor example:
$q=a(1-e)$
$\langle q\rangle \neq\langle a\rangle \cdot(1-\langle e\rangle)$


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## Conclusion:

 mean orbit described above is set of parameters which does not define a real orbit.
## Heliocentric vectorial elements

## - Angular momentum $h$

- Lenz vector $\boldsymbol{e}$

These vectors are perpendicular to each other. There are 5 independent components.


## Method description

- Take at least 7 orbits
- Pre-integrate orbits into one epoch
- Calculate heliocentric vectorial elements (h, e, E) for each orbit
- Average vectorial elements constraining $h$ and $e$ are perpendicular to each other and equation for energy is satisfied
- Calculate heliocentric orbital elements (q, e, i, $\omega, \Omega$ )
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- 500 clones of Neuschwanstein meteorite consistent with uncertainties of observed orbital parameters.
- Each meteorite was integrated for 5000 years.
- For each 10 year interval orbits were averaged using Jopek et al. method (blue line).
- Also best-fitting orbit was integrated (red line).



## D-criterion for integrated mean orbit and mean of integrated orbits



## 86 Quadrantids



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## Conclusions

- One should use proper method of mean orbit calculation.
- For integrated clones of Neuschwanstein mean orbit is close to most probable orbit.
- Quadarantids: calculation of mean orbit after integration of orbits gives spurious results.

References:
Jopek, T. J., Rudawska, R. \& Pretka-Ziomek, H. 2006, MNRAS 371, 1367

